**Identifying optimal precursor for two types of El Nino events based on CNOP with GFDL CM2p1 model**

**Abstract：**

**Key word:**

1. **Introduction**

El Nino-Southern Oscillation (ENSO) events are coupled ocean-atmosphere phenomena in the tropical Pacific and have received much attention for both their climatic and economic impacts(文献). More and more evidence shows that there are two different types of El Nino events occur in the tropical Pacific based on the interannual variation and spatial distributions of the sea surface temperature (SST): the eastern-Pacific (EP) El Nino and the central-Pacific (CP) El Nino(文献). Moreover, CP-El Nino have become more common during the late 20th century, especially after the 1990s(文献), which have increased difficulties of recognition and prediction of two types of El Nino events especially CP-El Nino. So how to identify the optimal precursor (OPR) for these two types of El Nino events is very significance for improving ENSO predictability.

Mu et al. (2003) proposed a novel concept of conditional non-linear optimal perturbation (CNOP), which is characterized by maximum nonlinear growth of the initial perturbation in a given condition. Many researches have used CNOP with climate models to study the OPRs for two types of El Nino. Xu (2014) used different cost functions to simulate OPRs for two types of El Nino by CNOP in Zebiak-Cane (ZC) model, the result shows that the OPRs can develop into EP-El Nino event and “mixed El Nino event” (Kug et al. 2009) except La Nina event. Duan et al. (2014) used optimal forcing vector approach to simulate EP-El Nino and CP-El Nino events as well as in ZC model, which can reduce the effects of model errors. However, both research implied the ZC model has no ability to simulate CP-EL Nino. So, if choosing a high dimension large scale ocean-atmosphere coupling model, whether can get a good simulation for two types of El Nino then identify OPRs.

Geophysical Fluid Dynamics Laboratory Climate Model (GFDL CM) is a complex global ocean-atmosphere coupling model with subsurface processes that can be used to study the predictability of ENSO events (Kug et al., 2009, 2011). It has been proven GFDL CM has the ability to give a good simulation for two types of El Nino events(文献). In this paper, GFDL CM is chosen to identify the OPRs of two types of El Nino events. In other words, to identify the OPRs of two types of El Nino events is to solve the CNOP of GFDL CM.

The need of integrations of adjoint model during solving CNOP, however, is limited in complicated operational models that do not have an adjoint model such as GFDL CM. To avoid this restriction, intelligent algorithms are used to solve CNOP in complicated ocean-atmosphere coupling model. (文献举例….). The dimension reduction methods are used in intelligent algorithms and the feature spaces are not change usually. However, GFDL CM has a very high dimension (high-resolution ocean and atmosphere model) of calculation and spends 4 hours for 1 year of integration. If the feature space is fixed when heuristic search, the convergence will be very slow. So, there is a need to accelerate convergence when using intelligent algorithms. In this paper, a dynamic feature space particle swarm optimization (DFPSO) is proposed to solve this problem.

1. **CNOP and Adaption function in GFDL CM**

Mu et al. proposed CNOP to study the predictability for weather and climate. Consider a nonlinear partial differential equation:

Where is the state vector, is a nonlinear operator, is the time. For the given initial field , the solution to Eq. (\*\*\*) is given by

where denotes the propagator of the nonlinear from time to time prediction time , GFDL CM will be used for nonlinear integration with given initial state. And the perturbation which makes the target function get the maximum with the condition of , as follows:

where is the nonlinear evolution of the initial perturbation , is the constraint of perturbation, and denotes the norm. In this paper, the perturbation is added on the sea temperature including sea surface temperature and subsurface temperature, the experiments in different type of El Nino event will have different range of perturbation. How to determine the range and will be introduced in Chapter (\*\*\*).

1. **Solving CNOP of GFDL CM with Dynamic feature space particle swarm optimization (DFSPSO)**

[图]

The most important indicator distinguishing two types of El-Nino events is the sea temperature including sea surface and sea subsurface. In GFDL CM, the ocean module is divided into 50 levels and the resolution of each level is . It is necessary to reduce the dimension at such a high resolution. PCA is adopted in this paper. Each iteration in dynamic feature space particle swarm optimization (DFSPSO) mainly contains two steps: 1) feature space iteration and 2) swarm advance. In step 1), because of the complexity of ocean-atmosphere grid data, which indicates that the number of grid is much larger than the number of samples, a Space-time conversion in PCA will be used when calculating feature space of samples. Moreover, some accelerate approximate calculation methods are also used here. In step 2), a dynamic advance step is used: using large steps to increase the search space at the beginning of the iteration for extend search scope, using small steps at the end of the iteration for accelerate convergence where searching near the optimal value.

* 1. ***Feature extraction***

A lot of studies have shown that any space can be projected onto a low-dimensional space and be represented approximately by the attractors(文献). According to this conclusion, PCA is an effective way to obtain the attractors of chosen samples.

The preprocessing of feature extraction is centralization of samples, that is,

After centralization, PCA will be used for dimension reduction. Supposed the sample Matrix is , where represents the number of grids and represents the number of samples. The principal components (PCs) can be acquired with eigenvalue Decomposition:

where represents the eigenvalues of with a descending order which is a diagonal matrix, and is composed by the eigenvectors referred by eigenvalues . In common sense, the number of grids is much large than the number of samples (), will be a very high dimensional matrix where the calculation process takes lots of time and effort. So, Space-time conversion will be used instead:

where . And the non-zero eigenvalues in and are the same. The relationship between and is as follows:

If the samples are independent each other or orthogonal, the number of eigenvalues for is , the number of eigenvalues for is , that is, only the first eigenvalues can be obtained. However, the actual application has little effect on the results, because usually only the first few most important eigenvalues will be focused on.

* 1. ***Dynamic feature space***

Usually the feature space is not changed in the algorithm. In the end of the iteration, however, the characteristics of optimal value cannot be well described. That is, the meaning of samples is usually only a quantitative accumulation. Most of the samples only have general characteristics but do not have the features we are looking for. The feature space calculated by such samples cannot grasp the characteristics of the optimal value at the end of the iteration, because most of the particles have evolved into particles with certain special features at the end of the iteration, and it is obviously not suitable to continue to use the original sample feature space.

The use of dynamic feature space in the iterative process is conducive to seizing the features searching for, and it can speed up the convergence and make the particle swarms reach the optimal value in advance. The dynamic feature space consists of sample replacement and re-calculate feature space. For speed up the calculation, a method of matrix perturbation is used.

***3.2.1 Sample replacement***

In each iteration in PSO algorithm, there are always some particles who have the larger adaption value according to adaption function, and they are regarded as better particles. Moreover, these better particles are “closer” to global optimum (or local optimum), so they are supposed to seize more features we want. Replace them in samples and re-calculate the feature space, the feature space will have characteristics closer to the optimal value. We choose to random replacement of the original samples.

***3.2.2 Matrix perturbation***

When the samples have changed, the feature space and eigenvalue decomposition should be re-calculated, which is a consumption of time and memory. So, the method of matrix perturbation is used, which is based on the difference in samples between iterations is small enough, that is,

where represents the samples of next iteration and represents the samples of last iteration. And if only a few particles replace into samples, will a very sparse matrix, and non-zero values will be small. suppose , represents the number of iterations, so the Eq. (\*\*\*) can be written as,

According to Rayleigh quotient(文献),

where is left eigenvector corresponding to eigenvalues, is right eigenvector corresponding to eigenvalues. Because is a hermite matrix, the left eigenvector is equal to right eigenvectors, and if has centralization process. So, expressed in the form of a matrix, Eq. (\*\*\*) is

According to Matrix perturbation and Rayleigh method(文献), the above formula needs some corrections,

After the above formula corrections, the and can be approximated as the eigenvalue and eigenvector of the samples.

* 1. ***Solve CNOP with DFSPSO algorithm***

The first thing to solve CNOP with DFSPSO algorithm in GFDL CM is generating samples for getting initial feature space, the number of samples is important for this algorithm, because it is equal to the dimension of particle according to the formula for projecting particles back to the original space

where represents the perturbation in GFDL CM ocean module, represents the particle in low dimension, can be obtained from Eq. (\*\*\*\*).

The dynamic feature space and dynamic step are added into the traditional PSO algorithm(文献), the advance iteration formula is as follows:

where represents the speed of th particle in th iteration, and represents the location of th particle in th iteration, represents the best location of th particle before th iteration, represents the best location of all particles before th iteration. is a random number in . Parameter , and will change dynamically as follows:

where represents the max step of iterations, represents the number of current iterations. In this way, and are monotonically decreasing functions, and is a monotonically increasing function. Using large steps to increase the search space at the beginning of the iteration for extend search scope, using small steps at the end of the iteration for accelerate convergence where searching near the optimal value.

During iterations, due to the randomness of the parameters, the particles may evolve into a larger value, if projecting them to the space of GFDL CM, they may present a larger sea temperature which does not meet the physical meaning. So, the constraint of sea temperature perturbation after swarm projecting back to original space is defined as follows:

where represents the location of grid in GFDL CM, represents the latitude of gird, represents the sea temperature perturbation of gird after the swarm projecting back to original space, represents the standard deviation of sea temperature perturbations of gird in samples. presents the constraint which can be determined according to the value of each sample calculated by this formula (Sorting in descending order, taking value of 10% of the ranking). Once the sea temperature perturbation after swarm projecting back to original space exceeds the constraint, the following formula can remap particle back into constraint:

where represents the swarm updated, represents the value of Eq. (\*\*\*) for this swarm.

1. **results and analysis**
   1. ***Environment and parameter settings***

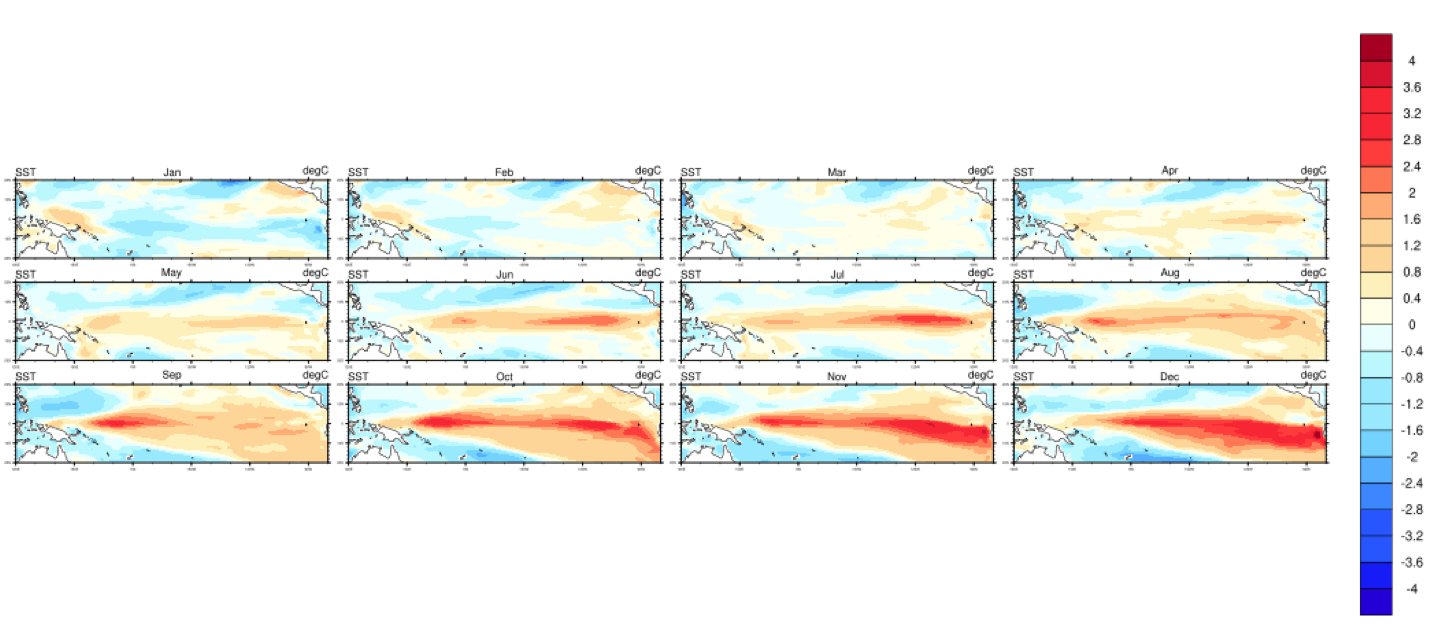
GFDL CM is a high-resolution ocean and atmosphere model which needs at least 45 cores to run, and DFSPSO is a kind of parallel particle swarm algorithm, so supercomputer (Tianhe2) is the best choice for this experiment. Due to limited resources, DFSPSO is set to run 10 particles in parallel at a time and the swarm count is 30 which means three rounds will complete the entire swarm. The DFSPSO parameters (, , ) are set like Chapter 3, is set to 20 based on experience. The samples are generated by GFDL CM itself which is the natural integration of the model for 350 years and discard its previous 50 years of unstable data which means the number of samples is 300. For different CNOP for two types of El-Nino, and are varied, the following chapter will give an introduction. As for solve OPR for two types of El-Nino, the “Spring prediction barrier” is an unavoidable question, so according to (文献), the January is chosen as start month, and December is chosen as predict month which can eliminate its impact as much as possible.

* 1. ***Optimal precursor for EP-El Nino***

Table (\*\*\*) shows the specific parameter settings of Solving CNOP of EP-El Nino in GFDL CM

|  |  |  |
| --- | --- | --- |
| Perturbation region | Perturbation level | Adaption function region |
| 30°N-30°S | 21 | Nino 3 area |

In this experiment, A normal year is chosen as the reference state, and the perturbation is added in January, Adaption function is calculated in December. Because the depth of the thermocline is the most important role in the Bjerknes positive feedback mechanism in EP-El Nino events, external tropical sea temperature signal is not very strong (文献), so the region of perturbation is determined between 30°N and 30°S and perturbation level is 21 including the SST and deeper under the sea surface (10m per level in GFDL CM). The step of iteration is 20, the CNOP is obtained as a result



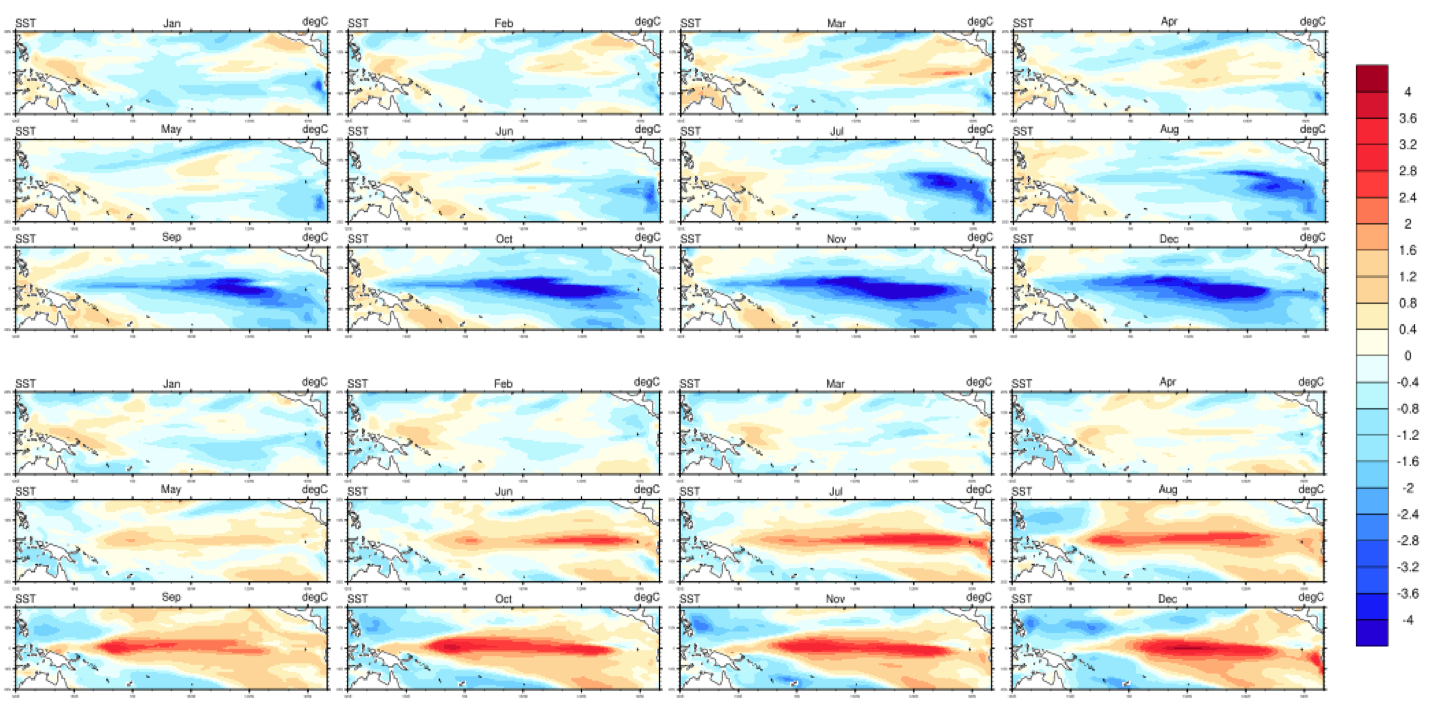
In Fig. (\*\*\*), the development of pattern is the global CNOP, which implies an EP-El Nino Event. In this pattern, the positive anomalies of SST first appear in the eastern Pacific Ocean, then move toward west until touching the west coast of the Pacific and move east, finally form a large value area in the Nino3 region.

* 1. ***Optimal precursor for CP-El Nino***

Table (\*\*\*) shows the specific parameter settings of Solving CNOP of CP-El Nino in GFDL CM

|  |  |  |
| --- | --- | --- |
| Perturbation region | Perturbation level | Adaption function region |
| 60°N-60°S | 11 | Nino 4 area |

In this experiment, the same normal year is chosen as the reference state, and the perturbation is added in January, Adaption function is calculated in December. Because external tropical sea temperature signal plays a more important role in CP-El Nino events than the depth of thermocline(文献), so the region of perturbation is determined between 60°N and 60°S and perturbation level is 11 including the SST and more shallow under the sea surface than in experiment of EP-El Nino. The step of iteration is 20, the global CNOP and local CNOP are obtained as a result

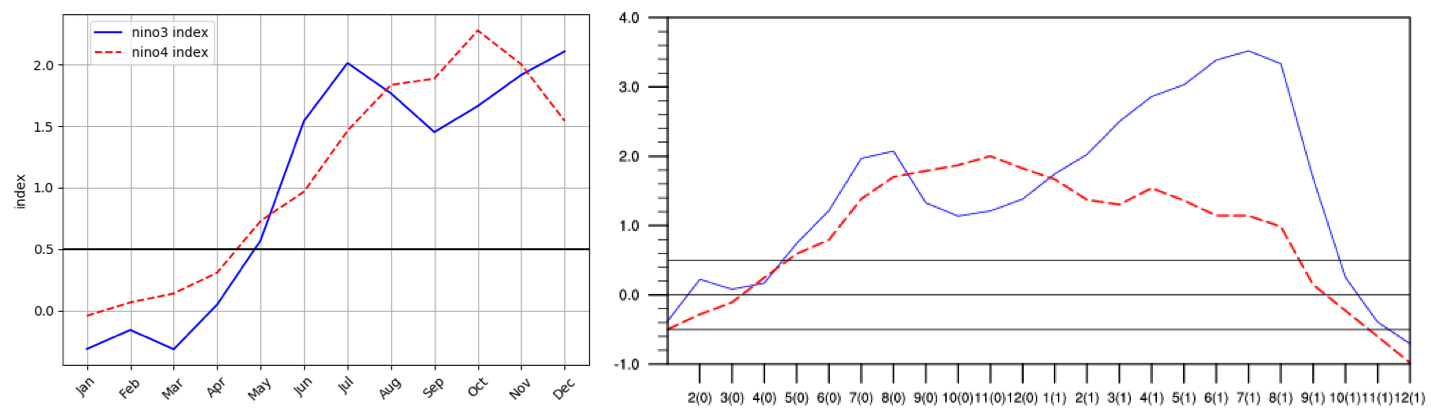


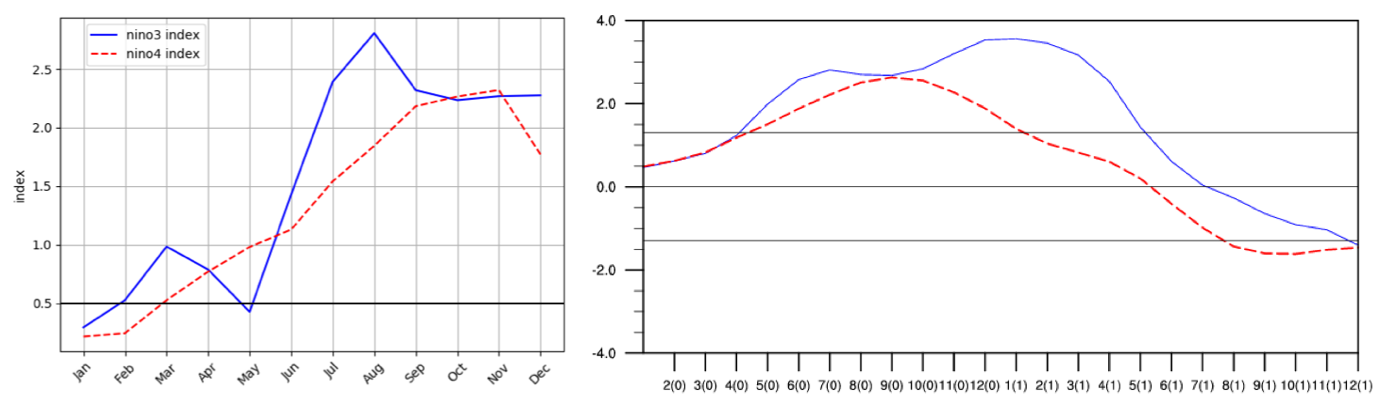
In Fig. (\*\*\*), the above is the global CNOP, which implies a La Nina Event, and the below is the local CNOP, which implies a CP-El Nino event. In the below development pattern, although the positive anomalies of SST are also distributed in the eastern Pacific Ocean, the positive SSTA in the Nino4 region also locally develops in the Nino 4 region, finally the SSTA large value area is in the Central Pacific Ocean.

* 1. ***Result analysis***

After repeating these two types of experiments respectively in some conditions: find two types of OPRs in different reference states and find them in the same reference states, two types of CNOP can be obtained successfully. The Magnitude and pattern of these two types of CNOP is almost stable. It is worth noting that, in the DFSPSO algorithm especially in CP-El Nino experiment, there exists two converge points: global CNOP and local CNOP, and the two points have different meanings in experiment. The reason for this phenomenon is that El Nino event and La Nina event always appear in pairs in nature which are both extreme climate phenomena (文献). The adaption value of La Nina is a little bigger than CP-El Nino when setting Nino 4 region for calculating adaption, but it is much smaller than EP-El Nino when setting Nino 3 region for calculating adaption. Moreover, in the CP-El Nino experiment, the DFSPSO algorithm can always find these two CNOP successfully, but in EP-El Nino experiment, the DFSPSO algorithm usually can only find the global CNOP which represents the EP-El Nino events. This is due to in the CP-El Nino experiment, the adaption values of La Nina event and CP-El Nino event are almost the same, in the EP-El Nino experiment, the adaption value of EP-El Nino event is much larger than the adaption value of La Nina event.

In order to verify the accuracy of the algorithm, Compare the development of Nino3 and Nino4 index of the two types of El Niño events obtained by DPSPSO algorithm with the two types of El Nino events resulting from the natural running of GFDL CM, and the results are as follows





It can be seen from the Fig. (\*\*\*) that the development trend of the two types of Nino index obtained by DFSPSO algorithm are almost the same as the two types of events in GFDL CM, Moreover, it is obvious that the SSTA and SSTA evolution in these two patterns look similar with the two types of events in GFDL CM, So the results of DFSPSO algorithm are reliable.

In addition, the DFSPSO algorithm spends less than 10 seconds averagely between iteration of the feature space matrix which is far less than eigenvalue decomposition directly. After a large number of experiments, the first four steps can almost find the global CNOP in DFSPSO. Fig. (\*\*\*) shows the adaption value development in the experiment, which indicates that: at the beginning of the iteration, the adaption value rapidly developed and reached the peak; during the middle of the iteration, the particle swarm increased the search range and searches for other fitness areas in the possible range which may cause a decline of adaption value; At the end of the iteration, the particle swarm quickly converge to the global CNOP. Nevertheless, the DFSPSO algorithm is a method of free of adjoint model, which makes it possible to calculate CNOP in high-resolution complex ocean-atmosphere coupling numerical models without adjoint models.



1. **Conclusions and future works**

This propose a DFSPSO algorithm for identifying OPR for two types of El Nino events, which solve CNOP in GFDL CM. DFSPSO algorithm includes the two main ideas: dynamic feature space and Iteratively calculating for feature space. Using these two methods can speed up algorithm convergence. It is worth nothing that the DFSPSO algorithm is a method which is free of adjoint models, and at the same time, it is stable and efficient and can get better results.

In addition, the DPFPSO algorithm can be run in parallel, which can increase the search space and the possibility of obtaining the optimal solution. In the future, a more complex adaption function will be designed which may be an effective combination of more physical quantities, and in this way, it is better to find the OPRs for two types of El Nino from the perspective of physical mechanism.